

MATH327: StatMech and Thermo

Thursday, 19 March 2026

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Something to consider

Quantum statistics should reproduce classical results when $\langle n_e \rangle \ll 1$
(so multiple particles are unlikely to occupy the same energy level).

What temperatures and chemical potentials correspond to this?

Recap
Classical Maxwell-Boltzmann $Z_g = \prod_{\lambda} \exp[\underbrace{-\beta(E_{\lambda} - \mu)}]$

Quantum statistics defines micro-states
by occupation numbers n_e
of energy level E_e

Boson and Fermions

Today

Bose-Einstein and Fermi-Dirac statistics
Classical limit

Photons (prep. for photon gas)

Grand-canonical part. func. for bosons

Sum $n_e = 0, 1, 2, \dots$ for every energy level E_e

Warm-up: Single E_0 with energy E_0

Micro-state energy $E_i = N_i E_0 = n_0 E_0$

$$Z_g = \sum_{n_0=0}^{\infty} \exp(-\beta(E_0 n_0 - \mu n_0)) = \sum_{n_0} [e^{-\beta(E_0 - \mu)}]^{n_0}$$

$$= \frac{1}{1 - e^{-\beta(E_0 - \mu)}}$$

Geometric series only converges

for $e^{-\beta(E_0 - \mu)} < 1$

$$\beta > 0 \rightarrow E_0 - \mu > 0 \rightarrow E_0 > \mu \quad \checkmark$$

$E_0 \geq 0$ $\mu = -T \left. \frac{\partial S}{\partial N} \right|_E < 0$

Generalize to E_ℓ with $\ell = 0, 1, 2, \dots, \mathcal{L}$

$\{n_\ell\}$ defines micro-states

$$Z_g = \sum_{n_0} \sum_{n_1} \dots \sum_{n_{\mathcal{L}}} \exp\left[-\beta \sum_{\ell} (E_\ell - \mu) n_\ell\right]$$

$$= \left(\sum_{n_0} e^{-\beta(E_0 - \mu)n_0} \right) \left(\sum_{n_1} e^{-\beta(E_1 - \mu)n_1} \right) \dots \left(\sum_{n_{\mathcal{L}}} e^{-\beta(E_{\mathcal{L}} - \mu)n_{\mathcal{L}}} \right)$$

$$= \prod_{\ell} \frac{1}{1 - e^{-\beta(E_\ell - \mu)}} \quad E_\ell > \mu$$

Bose-Einstein g.c. part. Func.

For fermions, only difference is $n_\ell \in \{0, 1\}$

Single $E_0 \rightarrow Z_g = \sum_{n_0=0}^1 e^{-\beta(E_0 - \mu)n_0} = 1 + e^{-\beta(E_0 - \mu)}$

General $E_\ell \rightarrow Z_g = \prod_{\ell} (1 + e^{-\beta(E_\ell - \mu)})$

Fermi-Dirac g.c. part. Func.

Collect results for $\Phi = -T \log Z_g$

$$\Phi_{MB} = -T \log \left[\prod_i \exp(e^{-\beta(E_i - \mu)}) \right] = -T \sum_i e^{-\beta(E_i - \mu)}$$

$$\Phi_{BE} = +T \sum_i \log(1 - e^{-\beta(E_i - \mu)})$$

$$\Phi_{FD} = -T \sum_i \log(1 + e^{-\beta(E_i - \mu)})$$

Classical MB should emerge from quantum

when $\langle N \rangle \ll$ number of accessible energy levels

$$p_i \sim e^{-\beta E_i} \rightarrow \text{more accessible } E_i$$

for small $\beta =$ high T

Compute $\langle N \rangle = -\frac{\partial}{\partial \mu} \Phi$ for all three and check high T

$$\langle N \rangle_{MB} = T \sum_i \frac{\partial}{\partial \mu} e^{-\beta(E_i - \mu)} = \sum_i \frac{1}{e^{\beta(E_i - \mu)}} = \sum_i \langle n_i \rangle_{MB}$$

$$\langle N \rangle_{BE} = -T \sum_i \frac{e^{-\beta(E_i - \mu)} \beta}{1 - e^{-\beta(E_i - \mu)}} = \sum_i \frac{1}{e^{\beta(E_i - \mu)} - 1} = \sum_i \langle n_i \rangle_{BE}$$

$$\langle N \rangle_{FD} = T \sum_i \frac{e^{-\beta(E_i - \mu)} \beta}{1 + e^{-\beta(E_i - \mu)}} = \sum_i \frac{1}{e^{\beta(E_i - \mu)} + 1} = \sum_i \langle n_i \rangle_{FD}$$

All occupation numbers ≥ 0 ✓

$$\langle n_i \rangle_{MB} = \frac{1}{e^{\beta(E_i - \mu)}}$$

$$\langle n_i \rangle_{BE} = \frac{1}{e^{\beta(E_i - \mu)} - 1}$$

$$\langle n_i \rangle_{FD} = \frac{1}{e^{\beta(E_i - \mu)} + 1}$$

All agree when $e^{\beta(E_i - \mu)} \gg 1 \rightarrow \langle n_i \rangle \ll 1$ ✓

looks like large $\beta =$ low T ???

Naive high-T limit $\beta \rightarrow 0$ with fixed $(E_i - \mu)$ gives

$$\langle n_i \rangle_{MB} \rightarrow 1 \quad \text{vs.} \quad \langle n_i \rangle_{FD} \rightarrow \frac{1}{2} \quad \text{vs.} \quad \langle n_i \rangle_{BE} \rightarrow \infty$$

Quantum effects important!

Reason: $\langle N \rangle$ also grows

True classical limit needs both

$T \rightarrow \infty$ (more accessible energy levels)

$-\mu \rightarrow \infty$ (keep $\langle N \rangle$ from growing)

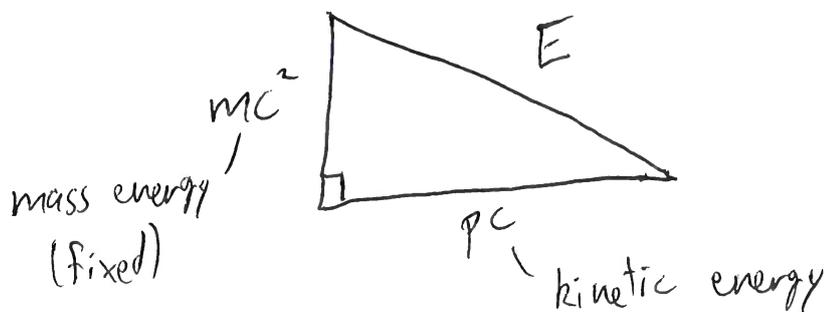
with $-\mu \gg T \gg E_i$ so $e^{\beta(E_i - \mu)} \gg 1$ ✓

Applications of quantum stats. in g.c. ensemble

need to specify energy levels E_i and energies E_i

In general $E^2 = (mc^2)^2 + (pc)^2$

$$p^2 = p_x^2 + p_y^2 + p_z^2$$



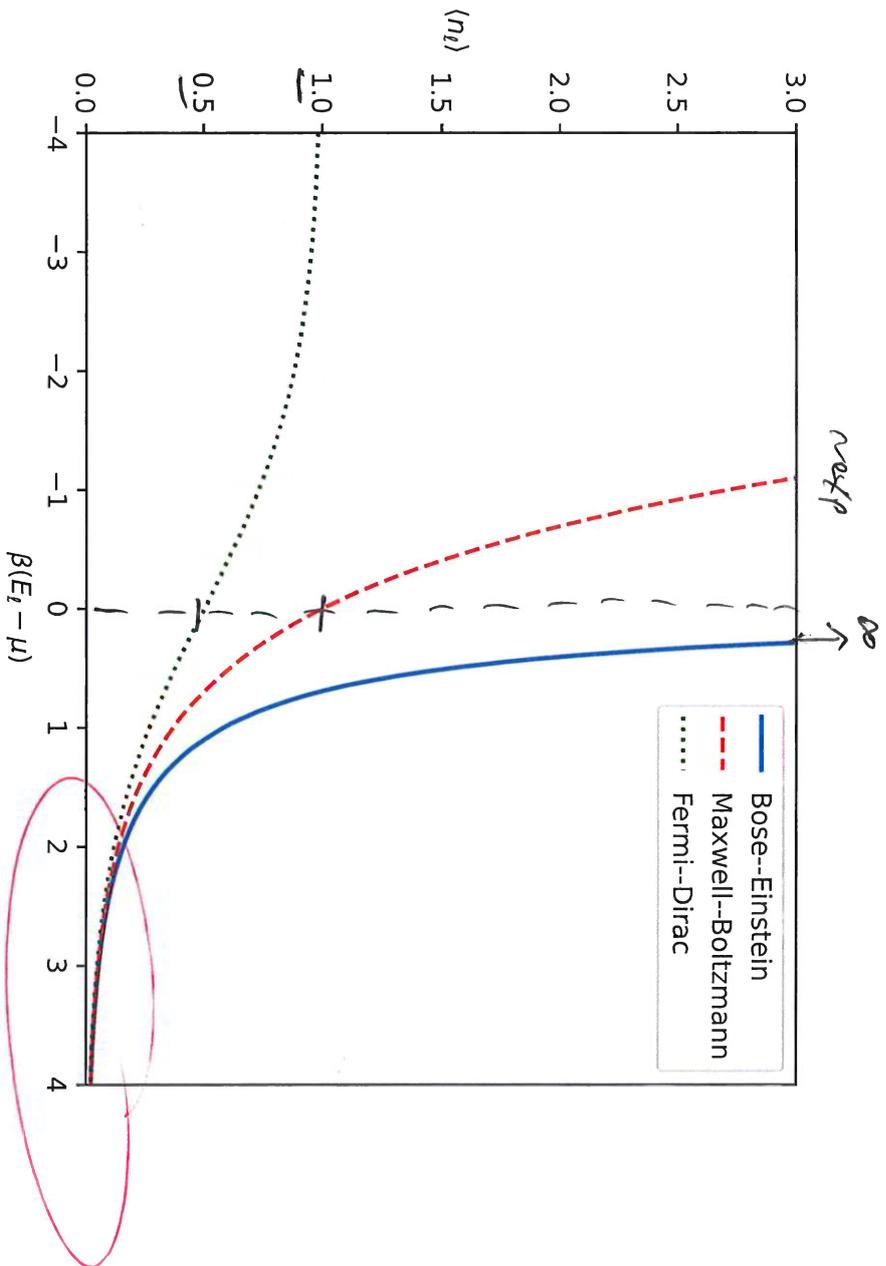
"Einstein's triangle"

We will consider two simpler limits

1) Ultra-relativistic $p \gg mc$ (including $m=0$)
 $E \approx pc$

2) Non-relativistic $p \ll mc$

$$E = mc^2 \sqrt{1 + \frac{p^2}{m^2 c^2}} = mc^2 + \frac{p^2}{2m} + \mathcal{O}\left(\left[\frac{pc}{mc^2}\right]^4\right)$$



Photons are bosons with $m=0 \rightarrow E_{ph} = pc$

Quanta of electromagnetic ~~waves~~ waves (light)

Constant speed of light $c = \frac{\lambda \omega}{2\pi}$

relates wavelength λ

and angular frequency $\omega = 2\pi F$

Volume $V=L^3$ quantize frequencies and energies

$L = k_i \left(\frac{\lambda}{2}\right) \rightarrow \lambda = \frac{2L}{k_i}$ "wavenumber" $k_{x,y,z} = 1, 2, \dots$

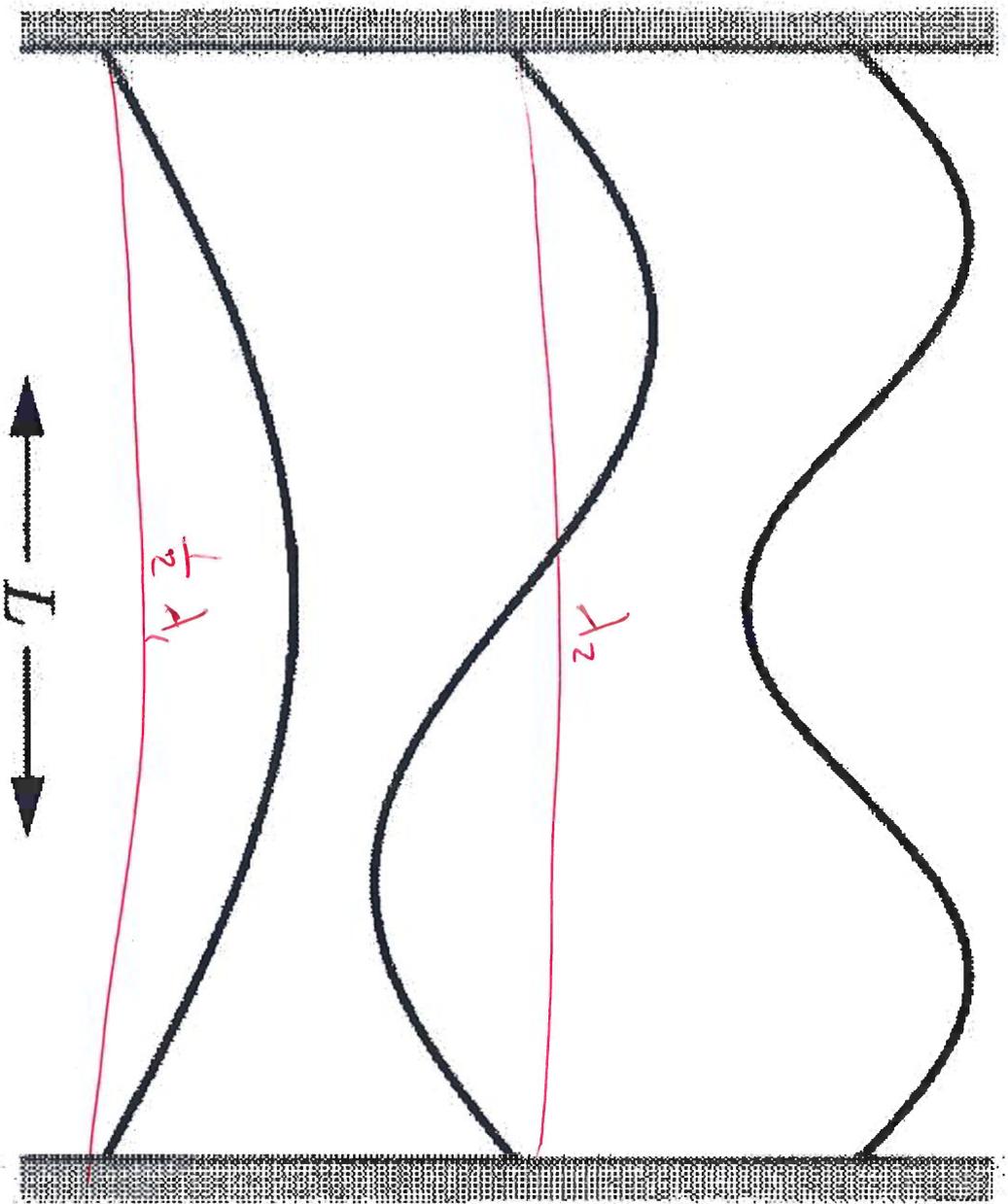
$$\omega = \frac{2\pi c}{\lambda} = c \frac{\pi}{L} k = c \left(\frac{p}{\hbar}\right) \quad p = \hbar \frac{\pi}{L} k$$
$$\sqrt{k_x^2 + k_y^2 + k_z^2}$$

$$\therefore E_{ph} = pc = \hbar \omega$$

Large $E \sim$ high freq. \sim short $\lambda = \frac{2\pi \hbar c}{E}$ "ultraviolet"

Small $E \sim$ low freq. \sim large λ "infrared"

For each $\vec{k} = (k_x, k_y, k_z)$, two degenerate energy levels
with different polarization

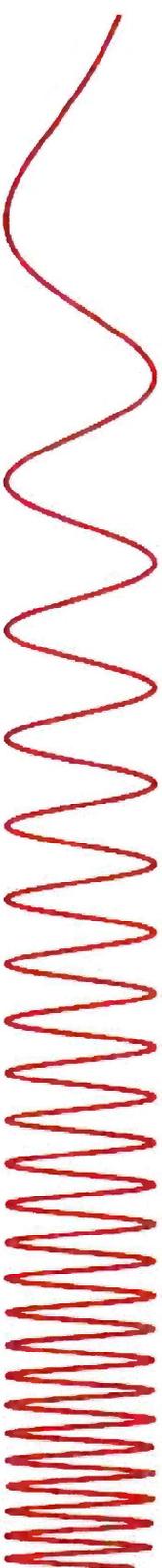


$$\lambda_1 = 2L$$

$$\lambda_2 = \frac{2L}{2}$$

$$\lambda_3 = \frac{2L}{3}$$

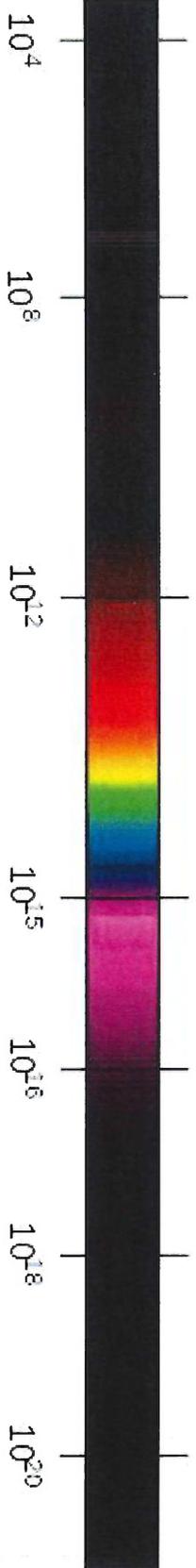
Penetrates Earth's Atmosphere?



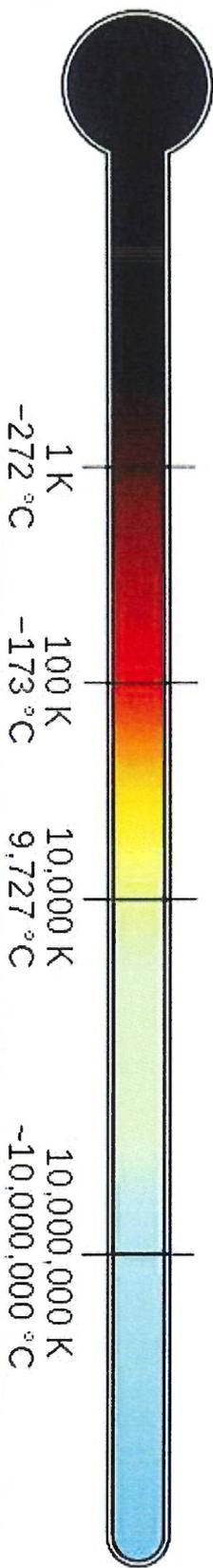
Radiation Type
Wavelength (m)
Approximate Scale of Wavelength

Radiation Type	Wavelength (m)	Approximate Scale of Wavelength
Radio	10^3	Buildings
Microwave	10^{-2}	Humans
Infrared	10^{-5}	Butterflies
Visible	0.5×10^{-6}	Needle Point Protozoans
Ultraviolet	10^{-8}	Molecules
X-ray	10^{-10}	Atoms
Gamma ray	10^{-12}	Atomic Nuclei

Frequency (Hz)



Temperature of objects at which this radiation is the most intense wavelength emitted



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